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1 **Laterally varying scattering properties from ambient**
2 **noise cross-correlations in the North Anatolian Fault**
3 **Zone**

4 Version: May 19, 2020

5 **Chantal van Dinther¹, Ludovic Margerin² & Michel Campillo¹**

6 ¹Université Grenoble Alpes, ISTERre, Grenoble, France

7 ²Institut de Recherche en Astrophysique et Planétologie, Université Paul Sabatier, Toulouse, France

Corresponding author: Chantal van Dinther, c.vandinther@gmail.com

8 **Abstract**

9 Intrinsic absorption and scattering properties provide us with information about
10 the physical state and heterogeneity of the Earth's crust. These properties are usually
11 obtained by observing the energy decay of naturally occurring earthquakes, leading to
12 sparse spatial sampling and therefore average scattering values over a large region. The
13 present study uses ambient noise cross-correlations to analyse the energy decay and scat-
14 tering properties over a part of the North Anatolian Fault (NAF; Turkey) from the con-
15 tinuous records of the 73 stations of the DANA temporary array in the frequency band
16 0.1 - 0.5 Hz. The region covered by the stations has rapidly varying geological charac-
17 teristics and is highly faulted around the northern strand of the NAF. We measured in
18 the noise correlations the space-time evolution of the energy of the coda waves. We first
19 perform measurements in separate sub-regions. The local scattering and attenuation prop-
20 erties are obtained by global optimization of a 2-D solution of the radiative transfer equa-
21 tion for surface waves. We found that the mean free path and attenuation coefficient are
22 considerably varying laterally with strong scattering observed in the region lying along
23 the northern strand of NAF. The optimization provides well-constrained values for the
24 scattering mean free path (ℓ) on the order of 10km in the fault region. The mean free
25 path is much larger ($>100\text{km}$) in the neighboring regions. We compare our global ob-
26 servations with simulations of scattered energy in a laterally variable scattering model
27 using 2-D radiative transfer. These simulations confirm the large contrast of heterogene-
28 ity between NAF and the surrounding crust and provide further constraints on the lat-
29 eral extent of NAF. When sources are located inside the fault zone, we find that energy
30 leakage controls the spatio-temporal distribution of coda wave energy in the medium.
31 This in turn suggests that lateral variations of scattering properties should be taken into
32 account in future monitoring studies.

33 **1 Introduction**

34 After the pioneering works of Aki (1969), it has been widely accepted that the coda
35 of seismic records is composed of waves scattered by heterogeneities in the lithosphere.
36 Aki and Chouet (1975) proposed to describe the energy decay in the coda as a combi-
37 nation of an algebraic and an exponential component. The latter is quantified by a fre-
38 quency dependent inverse coda quality factor Q_c^{-1} (Aki & Chouet, 1975) and varies with
39 the tectonic style of the region where it is measured (Singh & Herrmann, 1983). The en-

40 energy decay in the coda has since then been widely used to extract empirical information
 41 on the attenuation properties of the medium (Fehler & Sato, 2003). In spite of this suc-
 42 cess, it has become clear that precise information of the level of scattering cannot be ob-
 43 tained from the coda decay alone. Based on analytical solutions of the diffusion model
 44 in a half-space geometry, it has been proposed that Q_c^{-1} should be close to Q_i^{-1} , the in-
 45 trinsic quality factor of the crust (Aki & Chouet, 1975). However, Calvet and Margerin
 46 (2013) show that even in this simple geometry Q_c^{-1} and Q_i^{-1} only agree when scatter-
 47 ing is not too strongly anisotropic and at sufficiently large lapse-time. To properly as-
 48 sess the statistical properties of heterogeneities, and thereby extract detailed informa-
 49 tion on the Earth’s structure and composition, one needs to quantify and distinguish be-
 50 tween the intrinsic coefficient Q_i^{-1} and the scattering coefficient Q_{sc}^{-1} , both of them con-
 51 tributing to Q_c^{-1} through a relation that cannot be established in general.

52 One phenomenological method to investigate the relationship between the observed
 53 seismogram envelopes and the spectral structure of the random heterogeneity of the Earth
 54 is based on the scalar radiative transfer equation (RTE) (e.g. Hoshiaba, 1991, 1993, 1994;
 55 Margerin, Campillo, & Tiggelen, 1998; Sato, Nakahara, & Ohtake, 1997; Wu, 1985; Wu
 56 & Aki, 1988). In the case of a multiple-scattering medium there are two important pa-
 57 rameters that describe the heterogeneity: the scattering mean free path ℓ and the trans-
 58 port mean free path ℓ^* . ℓ , the reciprocal of the scattering coefficient, represents the av-
 59 erage distance between two scattering events. ℓ^* is the propagation distance required for
 60 a wave to lose memory of its initial direction. The first study taking multiple scatter-
 61 ing into account to estimate the relative contribution between scattering and intrinsic
 62 absorption was by Wu (1985). The work was based on a stationary transport equation.
 63 Later, time-dependent radiative transfer theory has been introduced to describe energy
 64 propagation in randomly inhomogeneous media (Sato, 1993a; Shang & Gao, 1988). The
 65 first Monte-Carlo simulations using radiative transfer for envelope synthesis were devel-
 66 oped in parallel (e.g. Abubakirov & Gusev, 1990; Gusev & Abubakirov, 1987). Hoshiaba,
 67 Sato, and Fehler (1991) and Fehler, Hoshiaba, Sato, and Obara (1992) analysed the en-
 68 tire S-seismogram envelopes (including the ballistic wave) to measure the ratio of scat-
 69 tering attenuation and intrinsic attenuation quantitatively. Their method, known as the
 70 multiple lapse-time window analysis (MLTWA), has lead to a large number of studies
 71 reporting regional values of the scattering properties (e.g. Mayeda, Koyanagi, Hoshiaba,
 72 Aki, and Zeng (1992) in Hawaii, Long Valley and Central California; Hoshiaba (1993) in

73 Japan; Jin, Mayeda, Adams, and Aki (1994) in Southern California; Carcolé and Sato
74 (2010) using the Hi-Net stations in Japan). Sato (2019) provides a comprehensive re-
75 view of scattering mean free paths measurements from rock sample to lithospheric scales.
76 The majority of these studies use earthquake records of local earthquakes with mag-
77 nitudes large enough for evaluating coda properties at large lapse times. This limits the
78 spatial coverage and therefore provides average or regional values of scattering proper-
79 ties instead of more specific local values.

80 Ambient noise cross-correlations (e.g. Shapiro & Campillo, 2004) offer an attrac-
81 tive alternative to study the attenuation parameters on a more local scale, particularly
82 in the lower frequency bands ($f < 1$ Hz). From scattering theory we know that we can
83 reconstruct the Green's function between two stations if we correlate the coda of an im-
84 pulse source, such as an earthquake (Campillo & Paul, 2003). Furthermore, it has been
85 shown that we can use the coda of correlations to also reconstruct the Green's functions
86 (Stehly, Campillo, Froment, & Weaver, 2008). This is a good indication that the coda
87 of correlations is analogous to the earthquake coda and contains valuable information
88 on the propagation properties.

89 Wegler and Sens-Schönfelder (2007) were the first to estimate the coda attenua-
90 tion Q_c^{-1} from ambient noise auto-correlations in Japan. More recently, Soergel et al.
91 (2020) mapped Q_c in the greater alpine region using ambient noise records in the 2.5-
92 20s period band and found large spatial variations of coda attenuation related to the ge-
93 ology. Hirose, Nakahara, and Nishimura (2019) demonstrated that cross-correlations func-
94 tions (CCF) can be used to derive scattering properties by comparison with active shot
95 data. In their study they derived average scattering properties at the Sakurajima vol-
96 cano in Japan by minimising the misfit between the energy densities measured in CCFs
97 and the predictions of scalar RTE. Motivated by these recent results, we use the coda
98 of ambient noise CCFs to detect and quantify possible lateral variations of scattering prop-
99 erties across the North Anatolian Fault Zone. Although the mapping of these spatial vari-
100 ations of attenuation across the fault is interesting in its own right, it may also have some
101 important implications for the accurate location of velocity changes based on coda wave
102 interferometry (CWI). Indeed, geological and geophysical studies (see e.g. Ben-Zion &
103 Sammis, 2003, for a review) suggest that the assumption of homogeneous scattering prop-
104 erties underlying most CWI tomographic studies may well break down near fault zones.

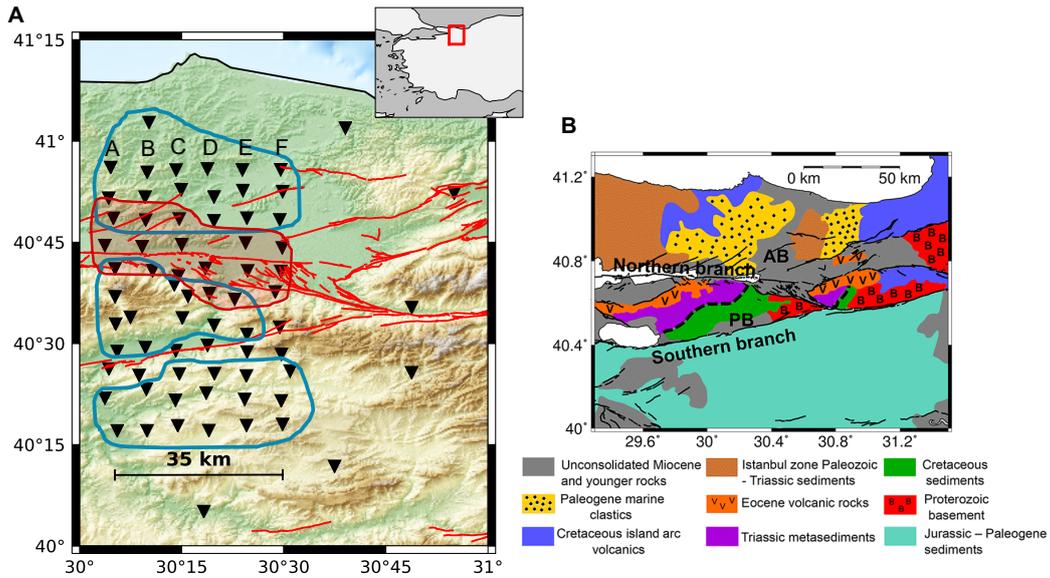


Figure 1. A) Map of the study region showing the DANA network (black triangles) and the main faults (red). The four sub-regions used in the data analysis are delimited by red lines for the fault zone and blue lines for the surrounding crust. The array is composed of 6 columns of stations, labeled a to f from west to east. Each column comprises 11 stations, numbered 1 to 11 from south to north. B) Geological map of the wider study region (from Taylor et al. (2019)). The Adapazari and Pamukova basin, are indicated by AB and PB, respectively.

105 Our study area is the North Anatolian Fault Zone (NAFZ; Turkey). This is a seis-
 106 mically active fault zone formed where the Anatolian block and the Eurasian continent
 107 meet. The ~1200km long dextral strike-slip fault has an east-west orientation. The re-
 108 gion around the Izmit rupture zone is the area we are considering in current work (see
 109 Figure 1). Here the NAFZ splays into a northern and southern strand; separating the
 110 study region into the Istanbul Zone in the north, the Armutlu Block in the centre and
 111 the Sakarya Zone in the south. The Istanbul zone consists of old and stiff continental
 112 material and a deep sedimentary basin, the Adapazari Basin (Şengör et al., 2005). The
 113 central block mainly comprises the Armutlu Mountains and has a small sedimentary basin
 114 in the southern part constrained by the southern strand of the NAF. The southern most
 115 part of our study region is the Sakarya Terrance. Similar to the central block it consists
 116 mainly of metamorphic rocks (Okay & Tüysüz, 1999; Yılmaz, Genç, Yiğitbaş, Bozcu, &
 117 Yılmaz, 1995). The northern strand of the NAF will be referred to as ‘fault zone’ (FZ)

118 in this study, because it has a wider faulted zone than the southern strand, due to the
 119 additional small normal faults.

120 The manuscript consists of two main parts. In the first part, we estimate the spa-
 121 tial variation of intrinsic attenuation Q_i^{-1} and scattering mean free path ℓ from ambi-
 122 ent noise cross-correlations (Section 3) using a data regionalization approach guided by
 123 geological considerations. In the second part we validate the inferred scattering prop-
 124 erties by direct comparison between the data and results of Monte Carlo simulations in
 125 media with laterally varying mean free path and intrinsic attenuation (Section 4). The
 126 next section is devoted to the description of the data set and signal processing.

127 2 Data processing and decay properties of coda waves

128 In this section, we describe the basic processing applied to the data. We show that
 129 the coda can be effectively reconstructed and perform an empirical analysis of its decay
 130 properties.

131 2.1 Data pre-processing

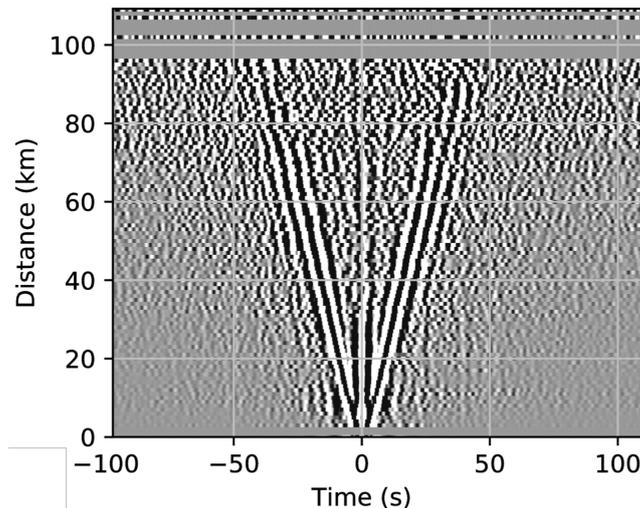


Figure 2. Cross-correlations between all pairs of N-components recorded at DANA in the 0.1-0.5Hz frequency band. The dominant arrivals show energy propagating at velocities that are typical of surface waves.

132 In this study we exploit ambient noise records from the Dense Array for North Ana-
133 tolia (DANA (2012); see Figure 1). This array consists of 73 three-component broadband
134 stations, 63 of which are forming a rectangle covering an area of approximately 35 by
135 70 km with an average inter-station distance of 7 km. The body wave reflection study
136 by Taylor, Rost, and Houseman (2016) and the surface wave tomography by Taylor et
137 al. (2019) provide convincing evidence that microseismic noise can be used successfully
138 at frequencies lower than 1Hz to reconstruct empirical Green’s functions from autocor-
139 relations and CCFs with the DANA array. To extract the direct Rayleigh wave and its
140 coda with sufficient signal-to-noise ratio, we computed CCFs using the complete 18 months
141 of continuous data which were recorded during the period between May 2012 and Oc-
142 tober 2013. The continuous data were first divided into one-hour segments. All compo-
143 nents were down-sampled to 25 Hz and corrected for the instrument response before re-
144 moving segments containing earthquakes of magnitude ≤ 2 as inferred from a local cat-
145 alog (Poyraz et al., 2015). A spectral whitening was applied to the data between 0.01
146 and 1 Hz, followed by 4th-order zero-phase Butterworth filtering in the 0.1-0.5Hz fre-
147 quency band. One-bit normalisation is the last step of the pre-processing and was ap-
148 plied to remove any remaining transient signal. We computed the full cross-correlation
149 tensor between all pairs of stations using 1h windows of pre-processed data. For each
150 station pair, the results from all windows were subsequently stacked to obtain the mean
151 CCFs over the full acquisition period.

152 Figure 2 shows an example of the resulting CCFs for the averaged horizontal com-
153 ponent pairs (for all 9 separate component pairs, see Figure 1 in Supplementary Mate-
154 rial). From the first arrivals we derive a velocity, which is approximately equal to 2.1 km/s,
155 indicating that the main energy pulse is most likely composed of surface waves. Hence-
156 forth, we will assume the coda to be mostly composed of scattered surface waves. For
157 the short inter-station distances, it is hard to differentiate between Rayleigh and Love
158 waves. This observation is confirmed by the surface wave tomography Taylor et al. (2019),
159 who show that the average velocities of Love and Rayleigh waves are very similar over
160 the study area in the 0.1-0.5Hz frequency band. To verify convergence of the calculated
161 CCFs we compare in Figure 3 the envelopes of CCFs derived from 12 months of stacked
162 data versus 18 months of stacked data. In the coda window the differences are negligi-
163 ble and we are thus confident that our CCFs have converged. The later part of the en-

164 velopes however, are different, showing that the noise levels are not the same across the
 165 study area.

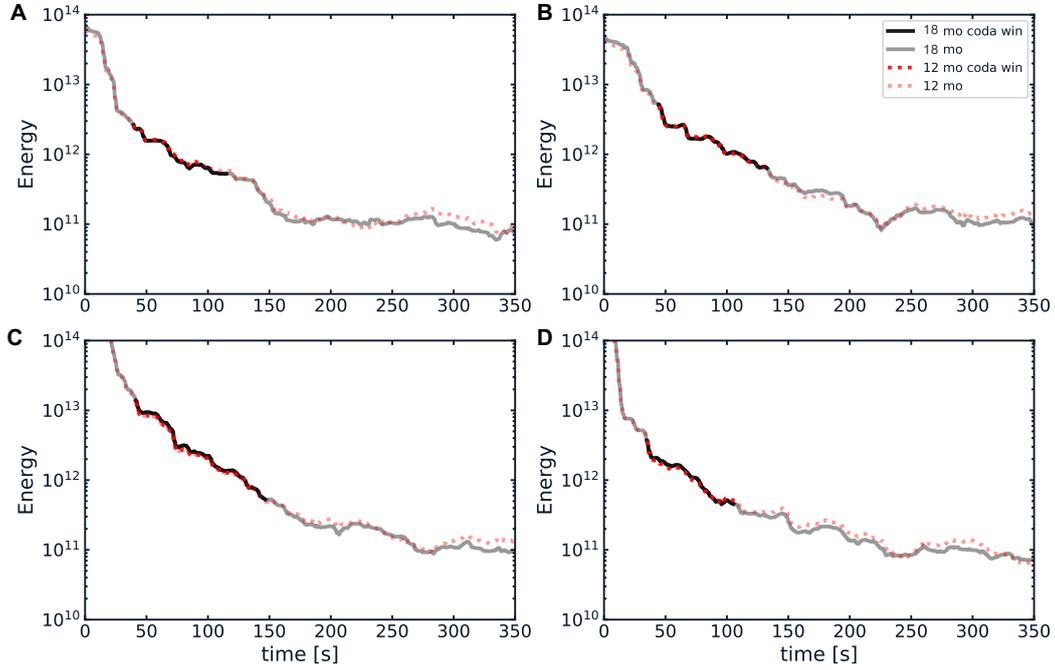


Figure 3. Envelopes of ambient noise cross-correlations obtained by stacking 12 months of data (red dotted line) versus 18 months of data (solid black line). The transparency of the curves is adjusted to delimit different portions of the data. The intense red and black part indicates the coda window, preceded by ballistic waves and followed by noise (high transparency). The four panels correspond to a representative sample of station pairs across the array.

166 **2.2 Measurement of Coda attenuation Q_c^{-1}**

A first estimate of seismic wave attenuation in Earth’s crust underneath the DANA array can be empirically obtained by measuring the energy decay of coda waves. Following Aki and Chouet (1975), we assume that the energy envelopes of noise CCFs obey the same algebro-exponential decay as earthquake data:

$$E(f, t) = S(f) \exp[-2\pi t f Q_c^{-1}(f)] t^{-\alpha} \quad (1)$$

167 where $E(f, t)$ is the mean-squared energy envelope at lapse time t around frequency f ,
 168 $S(f)$ is a frequency-dependent factor combining the virtual source magnitude and the
 169 site effect at the station and α is an exponent to be discussed below. $E(f, t)$ is obtained

170 by applying a smoothing window of 16 periods to the squared *CCF*, which is sufficient
 171 to remove rapid fluctuations of the envelope. To further improve the stability of the mea-
 172 surements, we subsequently average all four horizontal components of the energy envelopes
 173 of the cross-correlation tensor after a normalization of each term has been performed at
 174 a fixed lapse-time of 100s. The exponent α is a fixed parameter that depends on both
 175 the regime of scattering (from single scattering to diffusion) and the dominant wave type
 176 in the coda (body waves or surface waves). It typically varies between 1 and 2. Assum-
 177 ing that the ballistic and coda waves are mostly composed of surface waves, we choose
 178 $\alpha=1$. It is worth noting that in 2-D scattering media, the theoretical algebraic decay of
 179 energy is of the form t^{-1} in both the single-scattering and multiple-scattering regimes
 180 (Paasschens, 1997). Like in previous studies (Soergel et al., 2020), we estimate Q_c^{-1} di-
 181 rectly from the slope of the Log-Energy decay of the envelope against lapse time t with
 182 a linear least-squares method. We found that the assumption $\alpha = 1$ provides more sta-
 183 ble measurements of Q_c^{-1} than higher values ($\alpha = 1.5$ or $\alpha = 2$). We note that the
 184 quality of the coda reconstruction is not uniform over the network. In particular, sta-
 185 tions located in the south of the array show generally lower S/N ratios. To avoid under-
 186 estimation of the decay rate due to noise contamination, we decided to adapt the du-
 187 ration of the measurement window to the quality of the CCF. The coda typically starts
 188 25s after the ballistic arrival and ends when the signal to noise ratio (S/N) drops below
 189 5. We reject all CCFs for which (1) the coda duration is less than 75s or (2) the corre-
 190 lation coefficient of the linear regression R^2 is less than 0.75. This provides us with good
 191 quality estimates of Q_c^{-1} for the relatively early part of the coda. About $\sim 55\%$ of the
 192 constructed CCFs pass the selection criteria, leading to 1328 Q_c^{-1} measurements. We
 193 note that this amount of Q_c^{-1} measurements is for a maximum interstation distance of
 194 35 km (for details see next section).

195 **2.3 Mapping of lateral variations of Q_c^{-1}**

196 To facilitate the discussion of the results, the inter-station measurements of Q_c in
 197 the 0.1-0.5 Hz frequency band have been converted to 2-D maps. The study area is first
 198 discretized onto a grid of ($\sim 8.5\text{km}$ -by- 12km) pixels. At each pixel, we identify all the inter-
 199 station paths of total length smaller than 35km that propagate through and record the
 200 corresponding values of Q_c^{-1} . We then estimate the local value of Q_c^{-1} and its uncertainty
 201 by computing the arithmetic average of the recorded Q_c^{-1} values and their standard de-

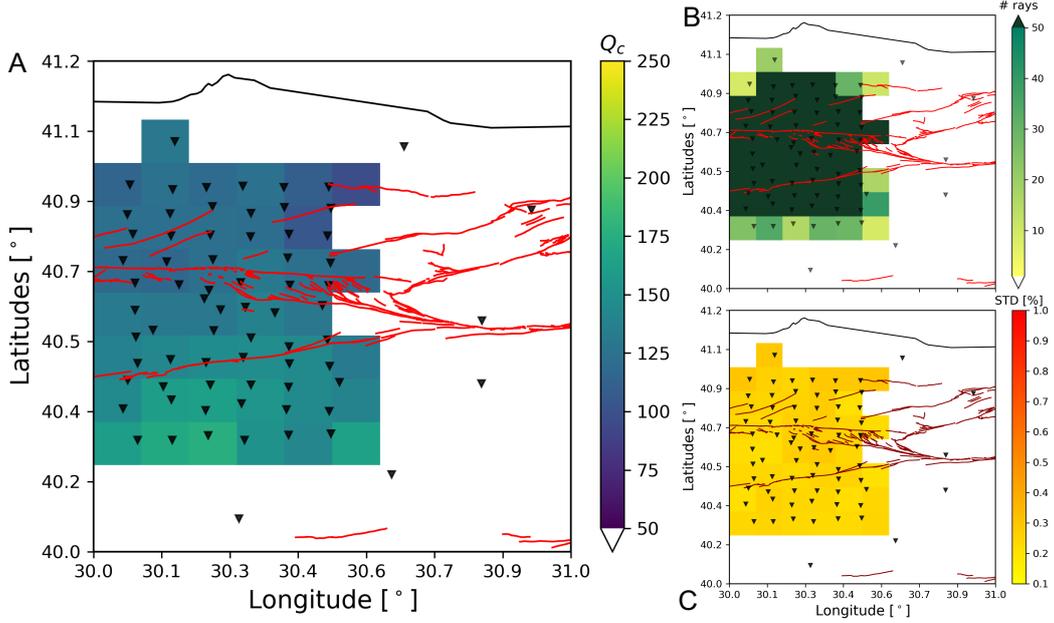


Figure 4. A) Q_c map for causal part of CCFs composed of the averaged horizontal components. It shows the arithmetic mean of Q_c for all rays that cross the cell with a minimum of 5 rays. B) ray coverage for averaged horizontal components C) standard deviation per cell in percentage of the measured Q_c .

202 variation. The choice of maximum path length (35 km) avoids mixing different propaga-
 203 tion regimes and provides a fine spatial resolution. The number of interstation paths per
 204 cell are displayed in Figure 4B. There are typically more than 50 rays crossing each pixel
 205 at the center of the array which ensures that the features shown on the map are reliable.
 206 In SM 2, we show that maps derived from both causal and acausal parts of different com-
 207 ponent combinations are very similar. Figure 4 illustrates that Q_c is relatively uniform
 208 over the entire study region. In particular, the fault zone does not distinguish itself from
 209 the surrounding crust. From this negative result, we may be tempted to conclude that
 210 the attenuation properties are uniform across DANA. This is not the case, however, as
 211 indicated by the direct observations of energy propagation presented below.

212 2.4 Laterally varying propagation properties

213 To give a direct image of the propagation properties of the region, and to detect
 214 potential local differences in properties, we have represented the energy distribution at
 215 different times deduced from the envelopes of the correlation functions. We considered

216 virtual sources located in two very different geological contexts. The first source is lo-
217 cated at the very south of the network (station DA01, Figure 5A) while the second is
218 located in the immediate vicinity of the northern branch of the fault (station DA07, Fig-
219 ure 5B). We measured the energy on non-normalised CCFs between these stations and
220 all others stations of the array. The energy is averaged for all horizontal components in
221 8 different time windows of 30 seconds, between 5 and 75 seconds. The energy snapshots
222 are presented in Figure 5 after temporal correction by the average Q_c . For a source south
223 of the fault system, Figure 5A does not show any spatial pattern after the initial flow
224 of energy from the source. The energy distribution exhibits a speckle-like behaviour, with-
225 out indication of any energy being concentrated at a particular location. On the con-
226 trary, when the source is close to the main fault, Figure 5B shows that the energy is not
227 spread uniformly but that is higher inside the fault zone, especially in the eastern side,
228 than outside the fault zone for lapse times between 25 and 65s. The absence of such a
229 pattern for the source in the south (Figure 5A) indicates that the energy concentration
230 observed in Figure 5B along the northern strand cannot be explained by a local effect
231 such as an amplification due to a shallow structure. One may notice that the region where
232 we find the energy concentration is close to a kink of the main fault where intense frac-
233 turing is expected (King, 1986) and observed (Figure 1). A dense fracture network would
234 result in a strong scattering strength for seismic waves. No clear energy concentration
235 is found for the scenario with a virtual source in the vicinity of the southern branch of
236 the fault zone. These direct observations suggest specific propagation characteristics
237 along the northern branch that are not revealed by Q_c mapping. In the following, we
238 rely on the Multiple Lapse Time Window Analysis (Fehler et al., 1992; Hirose et al., 2019)
239 to evaluate quantitatively the scattering and absorption properties in different sub-regions.

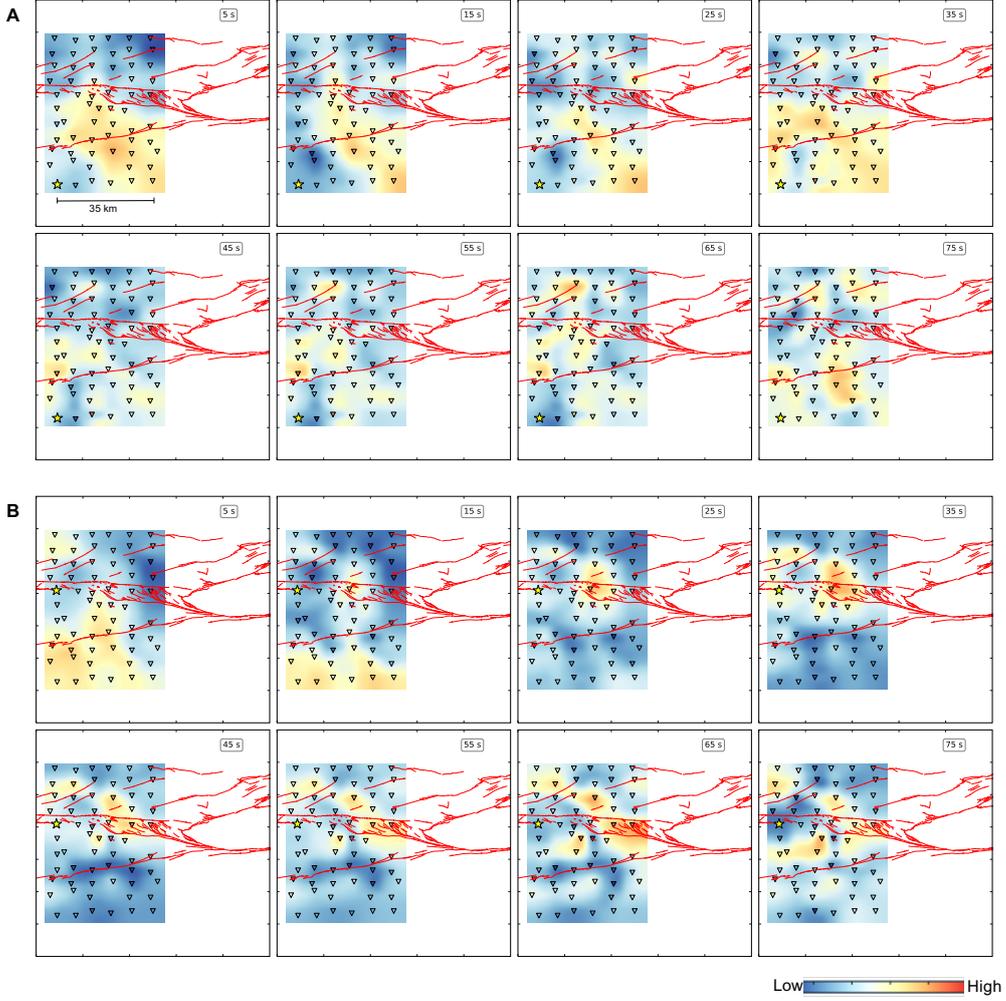


Figure 5. Snapshot of seismic energy distribution derived from ambient noise cross-correlations. The energy values have normalised for display purposes only. Assuming an average coda-Q of 150 in the study area and an average frequency of 0.3 Hz, the energy per time window is multiplied by $e^{2\pi ft/Q_c}$. A) The virtual source is located in the south at station DA01. B) The virtual source station DA07 (depicted by a yellow star) is located inside the fault zone. The color map indicates the energy intensity. Energy is entrapped in the eastern part of the fault zone up until 65s.

240 3 Mapping of attenuation properties

241 3.1 Transport model and inversion strategy

242 As stated in introduction, the attenuation properties of the medium as quantified
 243 by the intrinsic attenuation Q_i^{-1} and the scattering mean free path ℓ , are related to Q_c

244 via an unknown non-linear function. Although the previous section did not reveal sig-
 245 nificant variations of coda-Q over the region, it does not imply that the scattering prop-
 246 erties may not be variable. Since Q_c^{-1} especially relies on the assumption of a simple lin-
 247 ear decay of the Log-Energy in the time domain, it may be insensitive to finer details
 248 of the spatio-temporal distribution of energy, particularly at small spatial scales. To as-
 249 sess the possible contrasts in scattering properties between the fault and the surround-
 250 ing crust, we choose to split the study area into four sub-regions where we measure the
 251 scattering properties separately. We will refer to the sub-regions as North, ‘Fault Zone’
 252 (FZ), Centre and South, following a self-explanatory naming convention (see also Fig-
 253 ure 1). The definition of each sub-region results from a compromise between the follow-
 254 ing criteria. A sub-region should be: (1) small enough to ensure some homogeneity in
 255 the geology and therefore in the scattering properties; (2) large enough so that the aper-
 256 ture of the sub-array enables reasonable estimates of the scattering properties as further
 257 discussed below. We note that the selected stations in each sub-region lie inside the closed
 258 curves shown in Figure 1.

To estimate the local scattering parameters we perform a Multiple Lapse Time Win-
 dow Analysis (MLTWA) as originally proposed for earthquake data by Fehler et al. (1992)
 and recently applied to ambient noise CCFs by Hirose et al. (2019). In the case of the
 DANA data, we measure the total energy of the four horizontal components of the CC
 tensor in four 15-s long time windows starting at 5s, 25s, 50s and 75s after the ballis-
 tic arrival. The energy in a late time window of 15s duration starting at lapse-time $t=100$ s
 is used to normalize the measurements for the magnitude of the virtual sources and the
 site effects. The observed normalised energy densities (NEDs), E_{obs} , are then averaged
 in bins of width 2kms, in order to avoid bias towards one specific distance. The spatio-
 temporal distribution of NEDs in the four sub-regions is shown in Figure 6. It is appar-
 ent that the spatial energy decay in the first time window is much faster inside the FZ
 than outside, which suggests a contrast of attenuation properties between the FZ and
 its environment. To confirm this interpretation, we infer the values of the scattering mean
 free path ℓ and the intrinsic quality factor Q_i in each sub-region by comparing the ob-
 served NEDs to the predictions of a 2-D radiative transfer equation (RTE) applied to

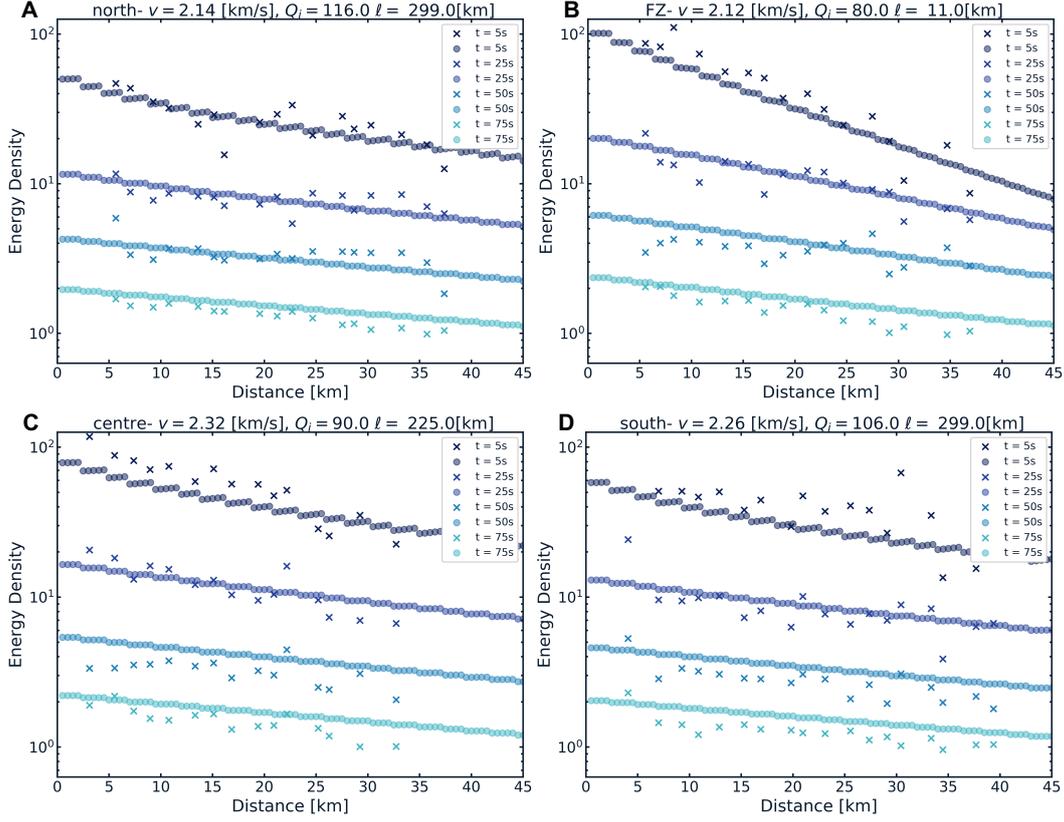


Figure 6. Normalised energy density (NED) in four sub-regions located inside the FZ (B) and outside the FZ (A, C, D). Crosses show the observed NED(‘x’) and circles (‘o’) the NEDs modelled by the 2-D RTE using the best fitting parameters indicated at the top of each plot. The color correspond to four time windows, with lighter blue colors for increasing lapse times (5s, 25s, 50s, and 75s after the ballistic waves).

the transport of surface waves (Paasschens, 1997; Sato, 1993b):

$$\begin{aligned}
 P(r, t) = & \frac{e^{-ct/l-\omega t/Q_i}}{2\pi r} \delta(ct - r) \\
 & + \frac{1}{2\pi lct} \left(1 - \frac{r^2}{t^2 c^2}\right)^{-1/2} \exp\left[l^{-1}(\sqrt{c^2 t^2 - r^2} - ct) - \omega t/Q_i\right] \mathcal{H}(ct - r)
 \end{aligned} \tag{2}$$

259 In Eq. (2), $P(r, t)$ represents the energy density at hypocentral distance r and lapse time
 260 t , for energy traveling at velocity c . The symbols $\delta(x)$ and $\mathcal{H}(x)$ represent, respectively,
 261 the Dirac delta function and the Heaviside step function. The first term on the RHS of
 262 Eq. (2) represents the direct wave contribution and is non zero only at $t = c/r$. The
 263 second term models the diffuse energy forming the coda for $t > c/r$. Note that the de-
 264 pendence of the intrinsic quality factor Q_i and the scattering mean free path ℓ (and in

265 turn of the energy density P) on the central frequency of the signal ($f = \omega/2\pi$) is im-
 266 plicit.

To quantify the agreement between the observed and simulated NEDs, we intro-
 duce the following misfit function:

$$SM = \sum_{i=1}^4 \sum_{j=1}^M \left[\log \left(\frac{E_{RTE}(t_i, j)}{E_{obs}(t_i, j)} \right) \right]^2 \quad (3)$$

267 where E_{RTE} denotes the energy density predicted by the 2-D transport model and the
 268 indices i, j refer respectively to the time window and the hypocentral distance bins. Note
 269 that the same normalization procedure is applied to the observed and modeled energy
 270 densities. In Eq. (3) the logarithm makes sure that all epicentral distances and lapse-
 271 time contribute equally to the misfit. To find the optimal value of mean free path and
 272 intrinsic attenuation, we perform a similar grid search as in Hirose et al. (2019) and Fehler
 273 et al. (1992). The search range for ℓ and Q_i^{-1} is the same for all sub-regions with ℓ vary-
 274 ing between 5 and 300 km with increments of 1 km and Q_i varying between 60 and 200
 275 with increments of 2. Before discussing the inversion results in the next section we briefly
 276 recall why separation of scattering and absorption properties is made possible by MLTWA.

277 The basic ideas were presented in Fehler et al. (1992) but we may revisit their ar-
 278 guments in the light of the sensitivity analysis of Mayor, Margerin, and Calvet (2014)
 279 which discuss the effect of local perturbations of attenuation properties. These authors
 280 show in particular the drastically different impact of scattering vs absorption on the seis-
 281 mogram energy envelopes. Scattering mostly affects the amplitude of ballistic waves and
 282 the early coda. If the scattering perturbation is located on the direct ray connecting the
 283 source and station, energy is removed from the direct waves and redistributed at later
 284 time in the coda. By contrast, a perturbation of absorption has a uniform impact on the
 285 energy envelop and affects the overall rate of decay of the energy in the time domain.
 286 For more complicated scenarios (e.g. scattering perturbation located off the direct ray),
 287 we refer to (Mayor et al., 2014). The fact that different time windows of the signal have
 288 quantitatively distinct sensitivities to elastic and anelastic perturbations is the key to
 289 MLTWA.

290 **3.2 Inversion results: absorption (Q_i^{-1}) and scattering (l).**

291 Figures 6 and 7 show the results of the optimization procedure for the scattering
 292 mean free path and the intrinsic absorption in each sub-region. The comparison of the

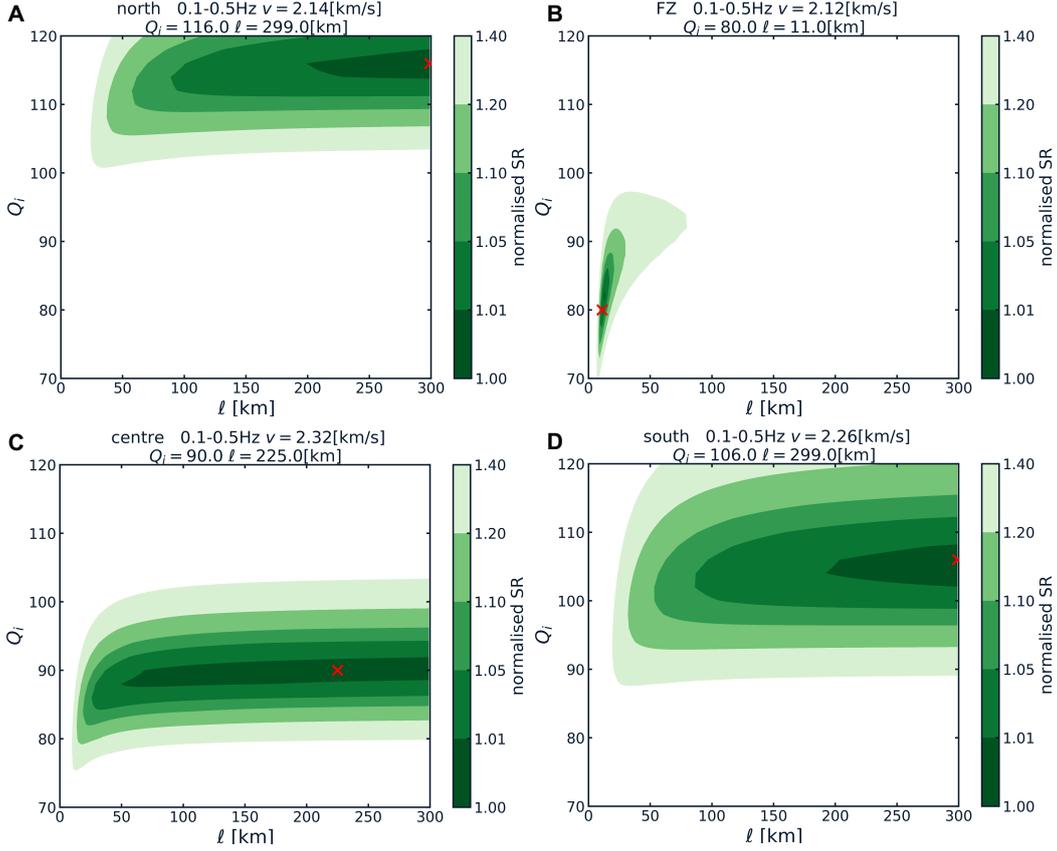


Figure 7. Resulting misfit and optimal values for the grid search in the 4 sub-regions. A), C) and D) for the regions outside the FZ, north, centre and south respectively; B) for inside the FZ. The green color indicates the values of normalised sum of the misfit (SM; as in Eq. 3), with the darkest color for the smallest misfit. The values are normalised w.r.t. the minimum SM. The best fitting values are indicated by the red crosses (‘X’).

293 predictions of the best-fitting model with the data indicates that a simple 2-D RTE is
 294 sufficient to capture the general spatio-temporal energy distribution across DANA, pro-
 295 vided that different attenuation parameters are employed in different sub-regions (see
 296 Figure 6). We first discuss the inversion results in the FZ. We note that the level curves
 297 of the misfit function in the $\ell - Q_i$ plane shown in Figure 7B indicate well-constrained
 298 values for ℓ of the order of 11 km and $Q_i \approx 80$, assuming a group velocity of 2.1 km/s
 299 deduced from the surface wave arrival times in the sub-region. The most outstanding
 300 observation which is reproduced by the RTE-model is the rapid spatial energy decay in-
 301 side the fault zone, mainly visible in the earliest time window of Figure 6. Compared to
 302 typical values reported in the literature (Sato, 2019), a scattering mean free path of the

303 order of 10 km is rather small. Yet it is still larger than the dominant wavelength of sur-
 304 face waves so that localization effects (see e.g. Hu, Strybulevych, Page, Skipetrov, & van
 305 Tiggelen, 2008) are probably negligible and the use of a transport model is legitimate.
 306 It is worth emphasizing that a scattering mean free path of 10km is not common in Earth's
 307 crust. Although comparisons with earthquake data are not straightforward due to the
 308 difference in frequency bands, comprehensive studies by Carcolé and Sato (2010) in Japan
 309 or Eulenfeld and Wegler (2017) in the United States suggest generally larger values of
 310 the order of 100kms or more. The level of absorption ($Q_i \approx 80$) is also found to be rather
 311 low but not exceptional for the sub-surface of the Earth. For the southern strand of the
 312 NAF we also performed the inversion, but no significant difference in neither the spa-
 313 tial decay rate nor the resulting ℓ and Q_i were observed compared to the neighboring
 314 central and southern sub-regions.

315 In the neighboring sub-regions, Q_i is well-constrained and the best-fitting values
 316 are slightly higher than inside the fault zone with $Q_i \approx 116, 90, 106$ for the north, cen-
 317 tre and south respectively. However, within a 20% error range Q_i shows similar values
 318 inside and outside the FZ and typically ranges between 80 and 120. From north to south,
 319 the surface wave velocity used in the simulations is 2.1km/s, 2.3km/s and 2.3 km/s, again
 320 derived from the arrival time of the direct waves. The values of Q_i are generally not con-
 321 sistent with the estimates of Q_c , as we would have expected the reversed trend with higher
 322 Q_i values in the south than in the north. This discrepancy may again be purely a con-
 323 sequence of the generally more complicated envelope shape of the data than the simple
 324 parametric form of the decay underlying Q_c measurements. The misfit contours of Figure7
 325 show that the mean free path is less well constrained outside the fault zone than inside.
 326 This is not surprising, as the mean free path ($\ell \geq 40\text{km}$) appears to be of the same size
 327 or larger than the largest linear dimension of the sub-network. The 20% error range in-
 328 dicates ℓ may range from 40km to over 300 km (300 km is the maximum value tested)
 329 in the normal crust. These values are typical of what is found worldwide (Carcolé & Sato,
 330 2010; Eulenfeld & Wegler, 2017; Sato, 2019).

331 **3.3 Effect of Velocity Model and Noise**

332 To asses the robustness of the optimization method and to better understand the
 333 poorer constraint on ℓ we explore the effect of the surface wave velocity and of the 'noise'
 334 in the data. For all four sub-regions different velocities are used, because they are de-

335 rived from the first arrival estimates per sub-zone. Similar to the Rayleigh wave tomog-
 336 graphy results of Taylor et al. (2019), we found the highest velocities in the centre and
 337 the lowest in the north. The actual values of the velocities, however, may be slightly dif-
 338 ferent from the ones we derived via the first arrival estimate, which affects the optimiza-
 339 tion. Although marginally better fits (smaller SMs) are obtained when using lower ve-
 340 locities in all sub-regions, there is no significant effect on the inverted scattering prop-
 341 erties (for more details see SM 3.1). Additionally, fluctuations in the NED measurements,
 342 or ‘noise’, affect the optimization and the resulting scattering properties as well. This
 343 is especially the case for zones with slow energy decay with distance in combination with
 344 a short array aperture. This aspect is also explored in greater details in the SM 3.2. The
 345 main findings from this section are that 1) Q_i is always well resolved because Q_i con-
 346 trols the typical energy ratio between the different time windows and 2) ℓ controls the
 347 slope of the first time window and may not always be well resolved due to the limited
 348 aperture of the array.

349 **4 Completing the cycle: comparing observations with energy trans-** 350 **port simulations**

351 To better understand the physical processes that play a role in the spatial and tem-
 352 poral energy propagation in a medium containing inhomogeneously distributed scatter-
 353 ers, we perform Monte-Carlo simulations of 2-D energy transport in a medium with pos-
 354 sibly variable ℓ and Q_i . Technical details of the numerical implementation are presented
 355 in Appendix A. Comparing the observations with the results from the simulations pro-
 356 vides us with insights about the scattering process in the NAFZ.

357 To facilitate the discussion, we compare two types of models: i) homogeneous mod-
 358 els, where the full space has the same scattering properties everywhere (further details
 359 may be found in the SM 4.1), and ii) inhomogeneous models, where space is divided into
 360 four different sub-regions including a fault zone, each with different scattering prop-
 361 erties. The motivations for this heterogeneous model are as follows (1) as discussed in sec-
 362 tion 3.2, the optimization is performed region by region, consequently there is no guar-
 363 antee of global agreement between the data and the model, and (2) as we have observed
 364 in Figure 6, the fit between the data and the model is not perfect, suggesting that there
 365 is room to improve the lateral variations in the model. To maximise the ability to com-
 366 pare the simulations with the observations, we use the same dimensions in the simula-

367 tions as in the study region. However, we halved the inter-station spacing to increase the
 368 number of NED measurements in the simulations. This is because we have in total more
 369 stations in the actual observations and therefore NEDs for most of the 2km bins.

370 **4.1 Constraints on the fault zone width**

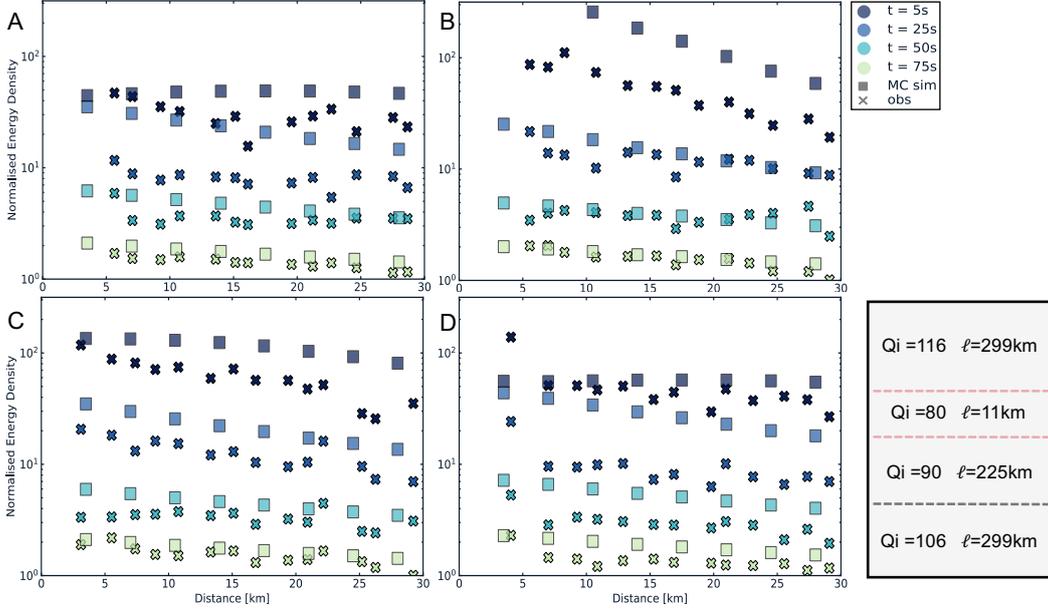


Figure 8. Comparison of spatio-temporal energy evolution between observations (crosses) and simulations (squares) for all sub-regions: A) north, B) FZ, C) centre and D) south. The colors indicate the different 15s-long time-windows, starting at 5s, 25s, 50s and 75s respectively. The results correspond to the initial heterogeneous model, based on the scattering properties derived from the regionalization approach and schematically shown in the lower right corner.

371 Two different configurations are used for the simulations in the case of inhomoge-
 372 neous models. The first configuration has east-west oriented receiver lines. They record
 373 the intensities for sources excited in the corresponding sub-region. An example of this
 374 configuration is shown in Figure 9. For these simulations we perform again the MLTW
 375 analysis and compare the results with the observations. Figure 8 shows that for all sub-
 376 zones and all time windows the NEDs are greater in the simulations than in the obser-
 377 vations, especially for the fault zone and the centre at early times. Taking simply the
 378 parameters derived in four homogeneous models and combining them in one heteroge-
 379 neous model seems insufficient to explain all observations in our study area. We have

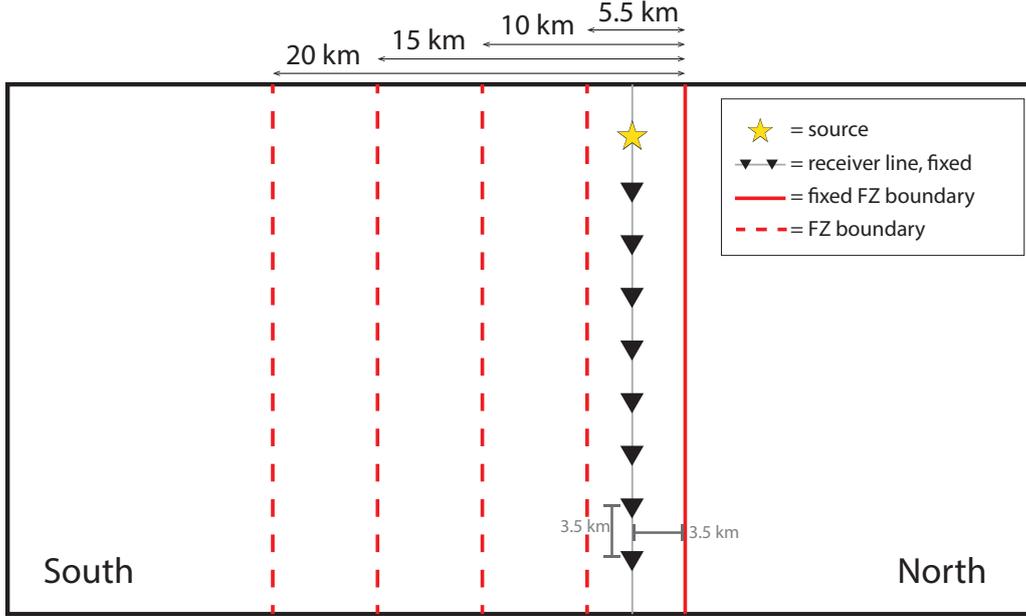


Figure 9. Example of the configuration for the simulations. Depending on the specific case, homogeneous or inhomogeneous, the model is divided in sub-regions where a source and an east-west receiver line is placed with receiver spacing 3.5km. Th specific configuration shown above is used for the inhomogeneous case with different fault zone widths. The different widths are realized by shifting the southern boundary (dashed red line). The fault zone boundary towards the north (right side) is fixed, as is the position of the receiver line w.r.t. the northern boundary and the source.

380 seen in the previous sections that (i) we constrain ℓ inside the FZ but it proved to be
 381 more difficult outside, and (ii) there is a clear difference in the decay rate with distance
 382 between inside and outside the fault zone (Figure 6). Consequently, the main focus of
 383 this section will be on explaining the observations in the fault zone. To simplify the anal-
 384 ysis we assume uniform scattering properties outside of the fault zone from here onward.

385 Figure 10 shows four panels with observation and simulation results inside the FZ.
 386 We first discuss the simple homogeneous models: panel A for a model space where Q_i
 387 =100 and $\ell = 10$ km, and panel B where $Q_i =100$ and $\ell = 150$ km. Both models do not
 388 match the observations to a desirable degree, especially in the first time window. For the
 389 model with strong scattering (A), the spatial energy decay in the first time window of
 390 the simulations is similar to the observations but the values themselves are too low. For

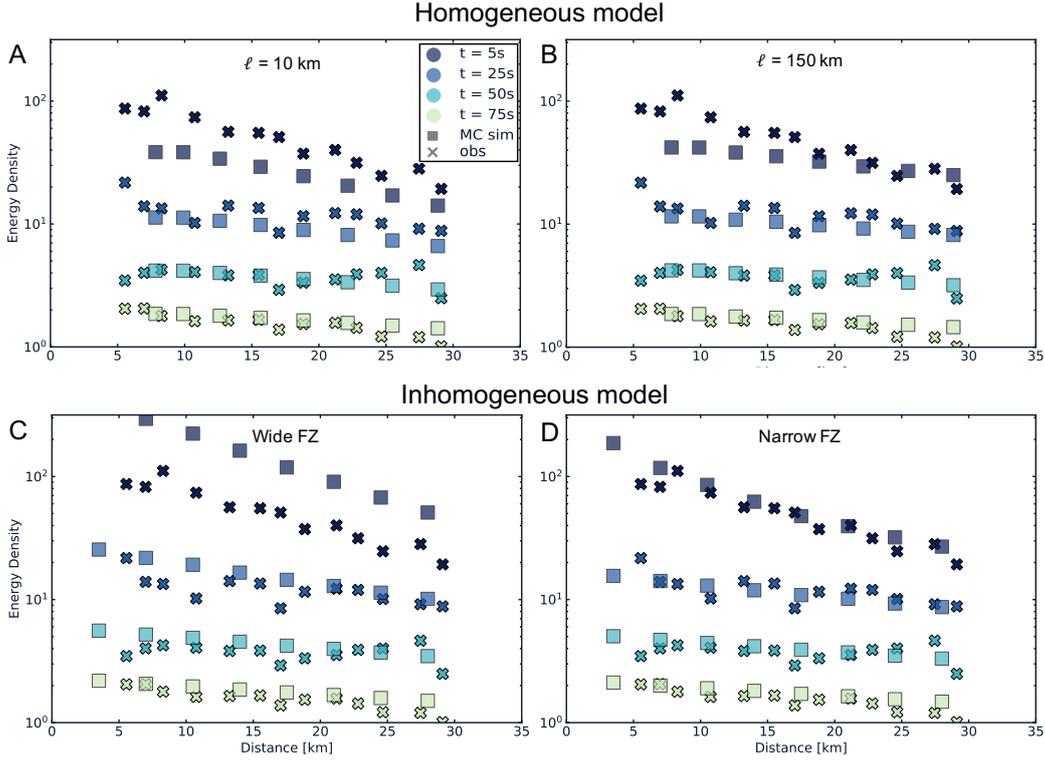


Figure 10. Comparison of spatio-temporal energy evolution between observations (crosses) and simulations (squares) for the FZ only. The upper two panels correspond to simulations in homogeneous models with A) $\ell=10$ km and $Q_i = 80$, B) $\ell=150$ km and $Q_i = 80$. The lower two panels are for a simple model with $Q_i = 80$ and $\ell=10$ km inside the FZ, and $Q_i = 106$ and $\ell=150$ km outside the FZ. C) for a wide fault zone, of 14 km and D) for a narrow FZ of 5.5 km. The colors indicate the different 15s wide time windows, starting at 5s, 25s, 50s and 75s respectively.

391 the weaker scattering model (B), it seems the other way around. The decay rate with
 392 distance at early times in the simulations is too slow, but the NED values are closer to
 393 the observations. Panel C and D show results for a simple heterogeneous model, with
 394 inside the FZ $Q_i = 80$ and $\ell = 10$ km, and outside the FZ uniform scattering properties
 395 for all sub-zones ($Q_i = 100$ and $\ell = 150$ km). The difference in the models between panel
 396 C and D is the width of the FZ, 14km for (C) and 5.5km for (D). Although the results
 397 for the wide FZ have the correct energy decay rate with distance, using a narrower zone
 398 with strong scattering yields an almost perfect match to the observations. This is in line
 399 with the findings of the tomography studies in the region (e.g. Kahraman et al., 2015;

400 Taylor et al., 2016), suggesting a fault zone that is perhaps not wider than 7km down
 401 to the mantle.

402 **4.2 The signature of a finite width scattering zone**

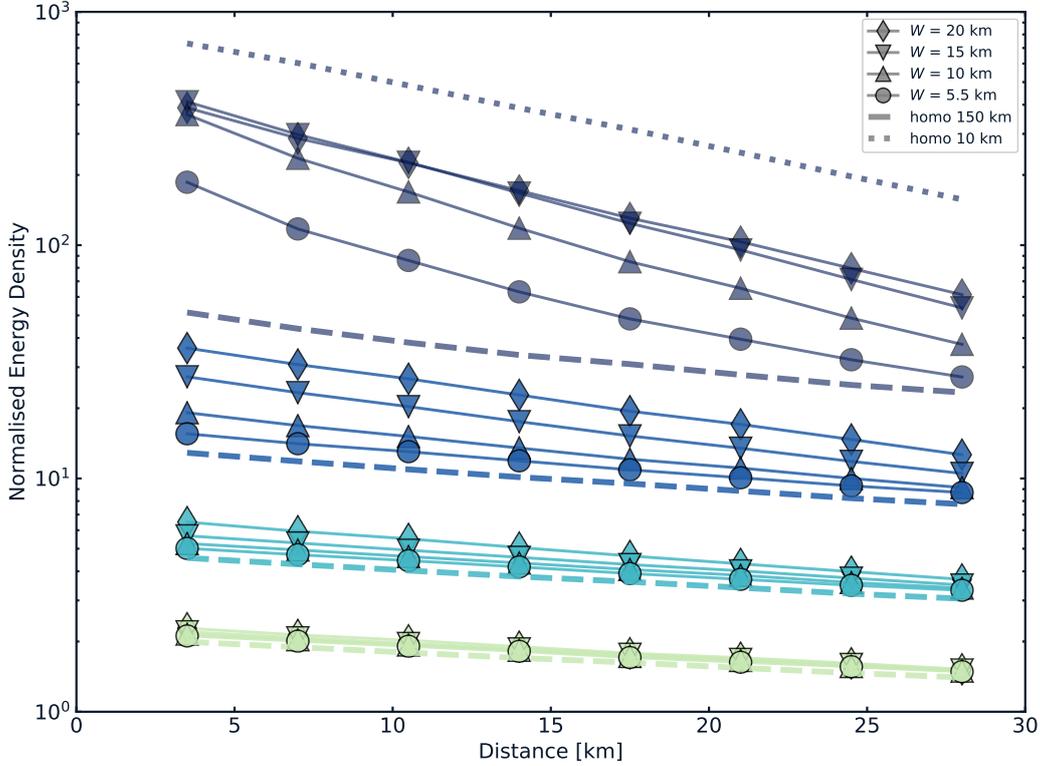


Figure 11. Comparison of spatio-temporal energy evolution for homogeneous and inhomogeneous simulations with different FZ width. The dashed and dotted lines show NEDs for homogeneous models with $\ell=150$ and 10 km, respectively. The NEDs of the inhomogeneous simulations are indicated by the continuous lines and markers, for different widths of the fault zone, W , as indicated in the legend. The colors indicate the different 15s wide time windows, starting at 5s, 25s, 50s and 75s respectively.

403 In the case of a structure with a band of material associated with high scattering
 404 taken between regions of low scattering, the energy decay with distance depends not only
 405 on the values of ℓ , as we saw in the numerical tests of the previous section, but also on
 406 the width of the band. Figure 11 shows the simulation results for fixed values of ℓ (10
 407 and 150km) and varying the band width W . We compare the behaviour of heterogeneous
 408 models with that of homogeneous models associated with the two values of ℓ 10 and 150km.

409 Note that for the heterogeneous models and the homogeneous model with $\ell = 150\text{km}$,
 410 the energies were normalized by their energy at 100s, as was done in the data analysis.
 411 A first observation is that in our time and distance configuration, the heterogeneous mod-
 412 els have very similar behaviors at 50s and almost identical at 75s. For the latter time
 413 the results are very similar to those of the homogeneous model with $\ell = 150\text{km}$. A first
 414 simple argument is the fact that the area of the strong scattering band becomes small
 415 in front of the total area covered by the scattered waves (Figure 9). The results pre-
 416 sented here have been normalized to 100s, to be consistent with the MLTWA analysis,
 417 but we have verified that the absolute long time energies for the heterogeneous and ho-
 418 mogeneous models with $\ell = 150\text{ km}$ are close (see Figure 16 in SM). Their relative dif-
 419 ference follows approximately a scaling in W/t , deriving from the same simple geomet-
 420 rical argument. We thus observe a convergence towards the solution of the external model
 421 with $\ell = 150\text{km}$ for long time in all cases. These results thus indicate that it is difficult
 422 to detect a heterogeneity from the observation of temporal decreases at long times, which
 423 explains why our characterization with Q_c was in vain.

424 The most important differences between the models are observed for the short times.
 425 Let us consider the case of the window starting 5 s after the arrival of the direct waves,
 426 which is critical in our comparison with the data. For this time window we have plot-
 427 ted in Figure 11 in dotted line the homogeneous model decay with $\ell = 10$, the value in
 428 the band representing the fault zone. Contrary to the other models and the data, the
 429 energies were not normalized with the 100s values for this homogeneous model. Based
 430 on our previous conclusions, the normalization was done with the values of the exter-
 431 nal homogeneous model ($\ell = 150\text{km}$) towards which the solutions of the heterogeneous
 432 models tend for long times. This normalization is carried out only for interpretation pur-
 433 poses, while approaching the data analysis conditions. Indeed, we know that the solu-
 434 tions converge towards the homogeneous case with $\ell = 150\text{km}$, but that they show dif-
 435 ferences at finite times (see Figure 16 in SM for non-normalized results). Since the band
 436 widths considered are small in relation to the distances travelled by the waves, and the
 437 sources are located near the band boundary, a significant portion of the energy escapes
 438 rapidly from the central band and will not be subject to intense scattering. Thus, a sig-
 439 nificant shift between the curve of the normalized homogeneous model and the hetero-
 440 geneous models is observed even from the short distances. For all widths, this shift in-
 441 creases with distance, which, it should be remembered, determines the absolute time of

442 the energy measurement window. This can be explained by the decreasing role played
 443 by the central band as the surface occupied by the diffracted waves grows, leading to the
 444 convergence of the heterogeneous models towards the homogeneous one with $\ell = 150\text{km}$
 445 for long times. We note that all heterogeneous models exhibit a decay with distance that
 446 is larger than the one of the homogeneous $\ell = 10\text{km}$ model.

447 The amplitude and decay of the normalized energies vary significantly with W . In
 448 the case where $W = 5.5\text{km}$, that is smaller than ℓ in the band, a significant shift is ob-
 449 served with the homogeneous model $\ell = 10\text{km}$ and the other models at the shortest dis-
 450 tance. The curve exhibits first a rapid decrease with distance then approaches the curve
 451 of the homogeneous model $\ell = 150\text{km}$, a convergence which is observed with the curves
 452 of the longer times. The amplitude is explained by the small size of the strongly diffract-
 453 ing band (5.5km) with respect to the radius or the region sampled by the waves. Even
 454 at the shortest distance, with the time window considered (here between 6.5 and 21.5s),
 455 the propagation of the waves is predominantly in the external environment. The posi-
 456 tion of the source in the band must also be taken into account when comparing ampli-
 457 tudes at short distances for the different geometries.

458 For the other values of W , the same evolution can be observed globally. For $t = 5\text{s}$,
 459 the curves gradually separate to approach the asymptote at distances increasing with
 460 W . The same behavior is observed for the other lapse-times. The evolution of the curves
 461 illustrates the transition between two limit models: strong scattering in the band for short
 462 times and weak for long times. Note that our configuration does not allow the first regime
 463 to be fully observed because the widths for the fault zone are small. The differences be-
 464 tween the heterogeneous models diminish rapidly with time or distance and the constraints
 465 on the model must be searched in a rather short time range but over a range of distances
 466 that covers the characteristic dimension/width of the heterogeneity. If we consider an
 467 observation at a fixed distance, e.g. 3 km , the amplitudes for W from 10 to 20 km are
 468 very similar for $t = 5\text{s}$ (the waves sample the same part of the model), much more sep-
 469 arated for 25s (at different convergence stages), then again very close for larger lapse-
 470 times (close to the asymptote). This example shows the need for a global vision of the
 471 behavior and a complete modelling. In particular, the effects of finite size of the diffract-
 472 ing band cannot be simply expressed by a single attenuation operator for each pair $(W,$
 473 $\ell)$ that would apply at all times.

4.3 Back-scattered Energy from the Fault Zone

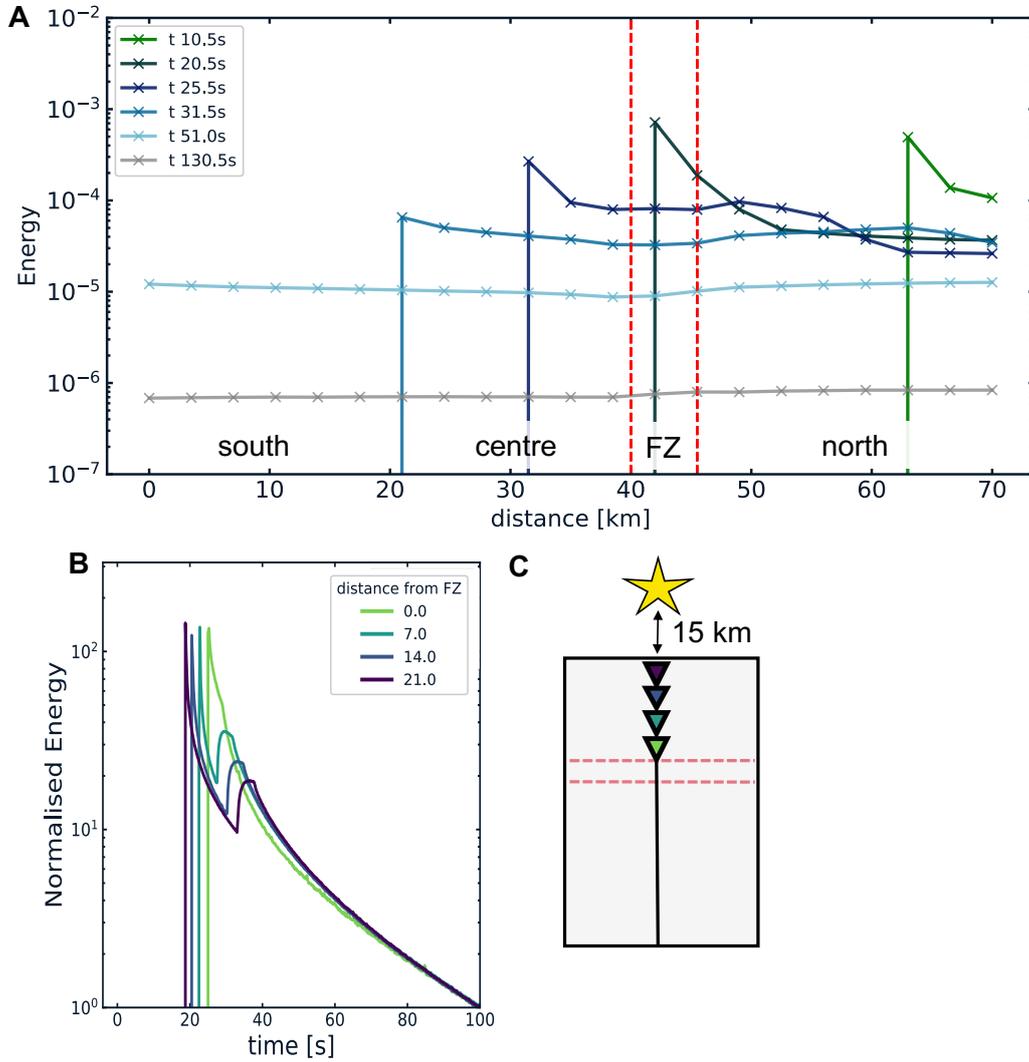


Figure 12. A) Snapshots of energy propagation, showing the effect of a strong scattering zone. The graphs show the energy density as function of distance (from right, north, to left, south) at different times. The red dashed lines indicate the fault zone. B) Normalised energy density versus time for four different receivers (7 km apart) in the same simulation. C) Simple heterogeneous model used for the simulation exhibiting a narrow fault zone, with $\ell/W = 2$, $\ell = 10$ km inside the FZ, $\ell' = 150$ km outside the FZ and $Q_i = 100$ everywhere. The receiver line is oriented north-south, in line with the source in the north, 15 km from the first receiver (yellow star). The color scale of the stations (C) and envelopes (B) indicates the distance between the station and the fault zone.

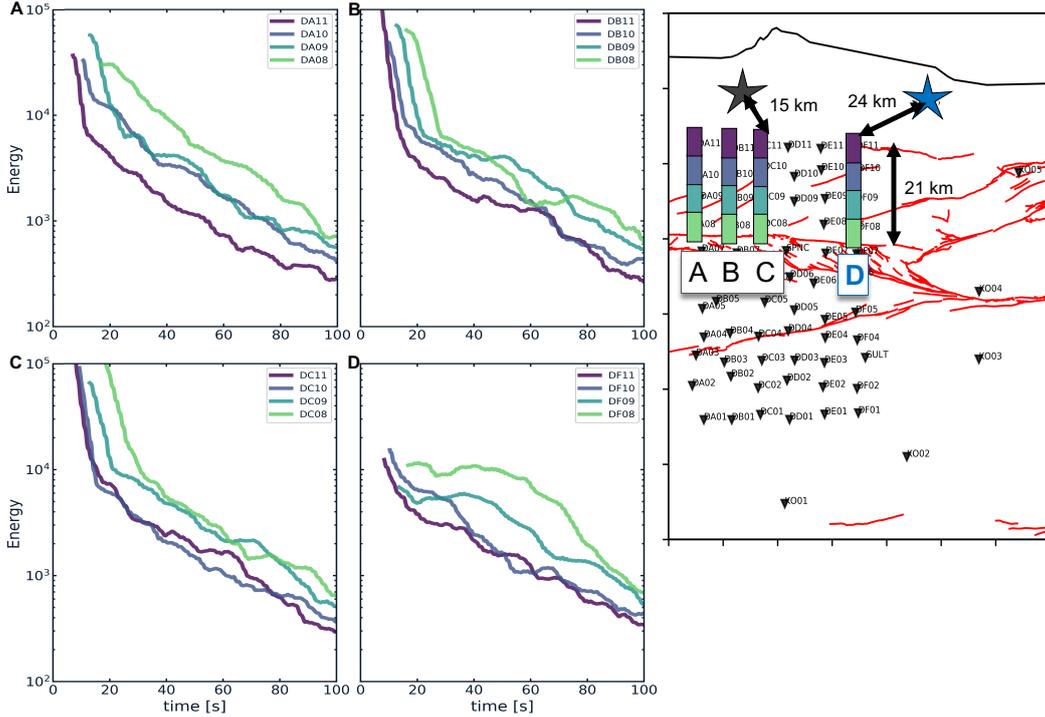


Figure 13. Observations of energy envelopes for station pairs oriented perpendicular to the fault zone. The same color scale as in Figure12B-C is used to indicate the distance between the stations and the fault zone. The map view shows sections A-D. For section A-C, the virtual source is the grey star (permanent station KO07). For section D, the virtual source is the blue star (permanent station KO06). The distance to the virtual sources are indicated on the map. Note: envelopes have been re-scaled with respect to the distance, d , between virtual source and station ($20 \times d$) to enhance the visibility of energy move-out.

475 In this section we explore the effects of a strong, narrow band of scattering on the
 476 energy evolution in space and time. The model we use is the one that seems most op-
 477 timal from the previous section (with a FZ width, W , of 5.5 km and $\ell = 10$ km inside,
 478 and $\ell = 150$ km outside the FZ). The configuration for these simulations is different than
 479 previously. The model space is again heterogeneous but the receiver line is north-south
 480 oriented. Snapshots of the resulting energy distribution, normalised energy envelopes as
 481 a function of lapse-time and the simulation configuration are shown in Figure 12A-C re-
 482 spectively. The most striking feature we observe on both the time-domain envelopes and
 483 the snapshots is a ‘bump’, which can be interpreted as back-scattered energy from the
 484 fault zone. When the ballistic energy reaches the fault zone, a front of diffuse energy prop-

485 agates back towards the direction of the source (Figure 12A). Figure 12B shows the same
486 feature, where for stations farther away from the FZ it is easier to distinguish between
487 ballistic energy and back-scattered energy. As we can see when the station is closest to
488 the FZ, both back-scattered energy and ballistic energy arrive at similar times. At a lapse
489 time of the order of 50s, we observe an inhomogeneous distribution of energy associated
490 with the lateral variations of scattering properties that progressively disappears at long
491 lapse-times (Figure 12A). If this phenomenon of strong back-scattering energy is also present
492 in the actual data, mixture of these energies can have profound implications for the coda
493 decay and scattering properties analysis. Unfortunately, identification of these bumps
494 in the actual CCFs proved to be more difficult than expected.

495 In the map view of Figure 13 we show four sections (A-D) perpendicular to the fault
496 zone, with a ‘virtual source’ roughly in line but as far as possible from the stations to
497 accommodate a move-out of this potential back-scattered energy. Sections A-C have per-
498 manent station KO07 (grey star) and section D has permanent station KO06 (blue star)
499 as their virtual source. The simulation (Figure 12B) discussed above is most similar to
500 section B of Figure 13. Additionally, the color codes for the envelopes are kept similar.
501 It is expected for the back-scattered energy to arrive later in the stations farther away
502 from the fault zone. Unlike the predictions from the simulations, we fail to observe back-
503 scattered energy in the data (Figure 13). The simple model with a strong narrow band
504 of scattering (Figure 12) seems to make predictions that are to discern in the data. This
505 suggest that the fault zone is far more complex than the simple conceptual model pro-
506 posed in this paper. Other medium properties such as the density and seismic velocity
507 may be significantly different inside the fault zone versus outside. This can cause a high
508 acoustic impedance contrast and therefore trap energy within the fault zone (Ben-Zion
509 & Sammis, 2003). In our simulations we have not taken other medium properties than
510 attenuation into account. It also seems reasonable to think that there are variations of
511 scattering strength along the strike of the fault. To investigate this possibility, we per-
512 form a virtual simulation of energy entrapment in the FZ using the real data.

5 Discussion

5.1 Fault zone trapped energy

As described in Section 2.4, energy seems trapped in the northern strand of the fault, for the case of a virtual source inside the fault zone (Figure 5B). For comparison, no energy entrapment is found for the scenario where the virtual source is in the southern most part of the study area (Figure 5A). The total energy distribution in the fault zone does not correlate with either the surface wave velocities (Taylor et al., 2019) or the earthquakes in the region (catalogue from <http://www.koeri.boun.edu.tr/sismo/2/earthquake-catalog/>). Additionally, Share, Allam, Ben-Zion, Lin, and Vernon (2019) showed energy entrapment in the San Jacinto fault zone using teleseismic and local earthquakes and a dense linear array along the fault zone. This favors the hypothesis of energy entrapment in the fault zone, and it may partly account for the remaining differences between the observations and simulations in Figure 10D. In the southern branch of the NAF we did not find any entrapped energy. This is in line with the findings mentioned in Section 3.2: the lack of significantly different scattering properties for this branch of the fault zone compared to the surrounding southern region. The absence of both concentrated energy and distinct scattering properties in the southern branch may be explained by the small damage zone, due to the lower activity of this fault (e.g. Aslan et al., 2019).

5.2 Fault zone complexity

We found a strong variation in scattering properties across the fault zone which promotes energy leakage. Evidently, the boundary between the fault zone and the surrounding medium is not purely of scattering properties. Our model, with simple back-scattering, provides a first order view but the actual structure is more complicated as observed in the remaining differences between the data and simulations. Effects of other physical quantities such as variations in seismic velocities and waves trapped in the fault zone have potentially a noticeable impact on the spatio-temporal energy evolution too. This corresponds to previous findings in the numerical study of Ben-Zion and Aki (1990), where they found waveform complexities and large amplitude amplifications due to head waves, surface waves and trapped modes in the presence of fault zone material heterogeneity. Additionally, section D in Figure 13 shows anomalies in the coda, compared to the codas of sections A-C. These anomalies in combination with the energy entrapment

544 towards the east of the fault zone, potentially also suggest along-strike variations. The
545 scattering mean free path may be smaller in certain parts of the fault than anticipated,
546 as well as the width, leading to a more transparent fault zone in certain areas. Further-
547 more, the more complex fault structure towards the east of the DANA array may leave
548 an imprint on the observed energy distribution.

549 **6 Conclusions**

550 This paper clearly demonstrates that a high scattering limited fault zone needs to
551 be included in a first order model for the DANA region. A narrow fault zone of ~ 5.5 km
552 width with a scattering mean free path of the order of 10 km, surrounded by a medium
553 with 150 km scattering mean free path, has been obtained by analysis of ambient noise
554 cross-correlations in the 0.1-0.5 Hz frequency band. We verified that our model repro-
555 duces a concentration of energy in the fault zone when the source is inside, as it is vis-
556 ible with the actual data (Figure 5B). This is clearly seen for long lapse times, that is
557 between 35 and 75s, in the simulations presented in SM Figure 17. Currently ambient
558 noise monitoring methods rely on the assumption that scattering properties are the same
559 across fault zones. Here, we have demonstrated that this is not the case. Therefore, lat-
560 erally varying scattering properties need to be taken into account in future monitoring
561 efforts.

562 From a methodological perspective we can conclude that one can use the coda of
563 correlations for attenuation studies. More importantly, this study reports for the first
564 time short scale variations in scattering properties. Although we know that scattering
565 properties vary between different types of crust, e.g. between volcanic and normal crust,
566 the derivation of the variations on such short scale has not been reported yet. Our re-
567 sults have direct implications for applications such as monitoring velocity changes in the
568 Earth's crust. Having an estimation of the scattering properties in one's study area can
569 have a significant impact on both the location and the quantity of temporal velocity changes
570 one finds. Future improvement of the model should be focused on incorporating lateral
571 variations along the fault zone, because the data revealed that there is even more com-
572 plexity than we describe by our model. A larger array, with an extension towards the
573 east, is necessary to capture the complexity.

574 **Acknowledgments**

575 The DANA (Dense Array for Northern Anatolia) array is part of the Faultlab project,
 576 a collaborative effort by the University of Leeds, Bogaziçi University Kandilli Observa-
 577 tory and Earthquake Research Institute (BU-KOERI) and Sakarya University. For the
 578 earth quake catalog we would like to thank Niyazi Turkeli. We also acknowledge the sup-
 579 port from the European Research Council (ERC) under the European Unions Horizon
 580 2020 research and innovation program (grant agreement No 742335, F-IMAGE).

581 **A Monte-Carlo simulations in laterally varying scattering media.**

582 We remind the reader that Monte-Carlo simulations designate a class of stochas-
 583 tic methods of solution of radiative transport equations. In this approach, energy is rep-
 584 resented by discrete seismic phonons who undergo a sequence of collisions upon which
 585 their direction of propagation changes according to the laws of scattering. In a medium
 586 with uniform scattering properties, the distance d between two collisions (also called free
 587 path length) is governed by an exponential probability law of the form $p(x < d < x +$
 588 $dx) = \ell^{-1} \exp(-x/\ell) dx$ with ℓ the scattering mean free path. For greater details, we
 589 refer the reader to the literature as summarized in Shearer and Earle (2004) and Sato,
 590 Fehler, and Maeda (2012). To simulate the transport of seismic energy in media exhibit-
 591 ing spatial variations of scattering and absorption, we have employed the so-called method
 592 of null or delta-collisions (Lux & Koblinger, 1991). This is an exact method of simula-
 593 tion that maps the transport process from a medium where attenuation properties vary
 594 spatially onto a medium where they are constant. Let us remark first that, in the frame-
 595 work of Monte-Carlo simulations, absorption may be treated as a scattering process that
 596 reduces the energy of a particle by a factor B , with B the local value of the albedo. Hence,
 597 to get a grasp on the method it is sufficient to treat the case of inhomogeneous scatter-
 598 ing properties, which is also most relevant to our applications.

Consider a 2-D transport equation with a spatially varying mean free path $\ell(\mathbf{r})$.
 The equation governing the transport of the energy is (e.g. Paasschens, 1997):

$$(\partial_t + c\mathbf{k} \cdot \nabla + \tau(\mathbf{r})^{-1})e(t, \mathbf{r}, \mathbf{k}) = \tau(\mathbf{r})^{-1} \int_{2\pi} p(\mathbf{k}, \mathbf{k}')e(t, \mathbf{r}, \mathbf{k}')dk' \quad (\text{A.1})$$

with $e(t, \mathbf{r}, \mathbf{k})$ the energy density flowing in direction \mathbf{k} (vector on the unit circle) at time
 t and position \mathbf{r} , $p(\mathbf{k}, \mathbf{k}')$ the scattering pattern governing the rate of transition from prop-
 agation direction \mathbf{k}' to propagation direction \mathbf{k} , c the seismic velocity, and $\tau = c\ell$ the

scattering mean free time. The integral on the left-hand side is over all the directions in the plane. We remark that if $e(t, \mathbf{r}, \mathbf{k})$ solves Eq. (A.1), it also solves the following equation where the effective scattering mean free time τ_e is *constant*:

$$(\partial_t + c\mathbf{k} \cdot \nabla + \tau_e^{-1})e(t, \mathbf{r}, \mathbf{k}) = \tau(\mathbf{r})^{-1} \int_{2\pi} p(\mathbf{k}, \mathbf{k}')e(t, \mathbf{r}, \mathbf{k})dk' + \tau_{nc}(\mathbf{r})^{-1} \int_{2\pi} \delta(\mathbf{k}, \mathbf{k}')e(t, \mathbf{r}, \mathbf{k})dk', \quad (\text{A.2})$$

599 where $\delta(\mathbf{k}, \mathbf{k}')$ is the delta function on the unit circle, $\tau_{nc}(\mathbf{r})$ is the scattering mean free
 600 time for the null or delta scattering events and $\tau_e^{-1} = \tau(\mathbf{r})^{-1} + \tau_{nc}(\mathbf{r})^{-1}$. In effect, all
 601 that we have done is to add the same term $\tau_{nc}(\mathbf{r})^{-1}e(t, \mathbf{r}, \mathbf{k})$ on both sides of Eq. (A.1).
 602 But the benefit is in fact immense because in the fictitious medium with null collisions,
 603 the scattering mean free path is constant, so that the most basic implementations ap-
 604 ply straightforwardly. Note that there is a certain degree of arbitrariness in the defini-
 605 tion of τ_e . Assuming that τ is bounded in the domain of interest, we may simply take
 606 $\tau_e^{-1} = \sup \tau(\mathbf{r})^{-1}$.

In practice, these ideas may be very easily implemented as follows. The transport process is simulated in a medium where the mean free path $l_e = c\tau_e$ is constant, so that the simple exponential probability law described above still applies. At each collision point \mathbf{r}_c , we must decide whether a true or a fictitious scattering event occurs. To carry out this task, we select a uniformly distributed random number in the interval $]0, 1[$ and compare it to the local scattering conversion rate. If:

$$\epsilon < \tau(\mathbf{r}_c)^{-1} / \tau_e^{-1} \quad (\text{A.3})$$

607 then a genuine scattering occurs and we proceed as usual with the selection of a new prop-
 608 agation direction. Otherwise, we have a null collision and the propagation direction is
 609 unchanged up to the the next collision. The main computational overload comes from
 610 the fact that we are in fact simulating more scattering events than actually occur in the
 611 real world. But this is largely counter-balanced by the considerable simplification of the
 612 free path length selection as well as the absence of complicated geometrical tracking of
 613 the particle.

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